

EXODUS: A mission to explore exoplanet evolution through understanding atmospheric escape

Citlali Bruce Rosete^{1*}, Mireia Leon Dasi², Marius Anger³, Paula Benitez Sesmilo⁴, Mark R. Boyd⁵, Frederik Dall'Omo⁶, Simone Filomeno^{7,8,9}, Jan-Vincent Harre¹⁰, Kim Angelique Kahle¹¹, Isabel Pitz¹², and Vito Saggese¹³

¹University of Luxembourg, Luxembourg; ²Observatoire de Paris, Université PSL, France; ³Aalto University, Finland; ⁴Polytechnic University of Catalonia, Spain; ⁵Imperial College London, United Kingdom; ⁶University of Stuttgart, Germany; ⁷Osservatorio Astronomico di Roma, Italy; ⁸Università di Roma Tor Vergata, Italy; ⁹Sapienza Università di Roma, Italy; ¹⁰German Aerospace Centre (DLR), Germany; ¹¹Max Planck Institute for Astronomy, Germany; ¹²Technische Universität Berlin, Germany; ¹³Università di Napoli Federico II, Italy

Introduction

Understanding the variety of system architectures and formation histories of planetary systems remains a major challenge. Current detection methods are strongly biased towards short-period bodies, leaving a gap in the exoplanet population demographics [1], [2].

Scientific questions:

1. How does **atmospheric escape** shape the evolution of long orbital period exoplanets?
2. What proportion of the exoplanet population do planets with **long orbital periods represent**?
3. How does the **Solar System architecture** compare to that of exoplanet systems?

Exodus is a mission to study the largely unexplored population of sub-Neptune to Jupiter-sized exoplanets with orbital periods greater than 100 days, through new detections and characterisation of atmospheric escape [3-7].

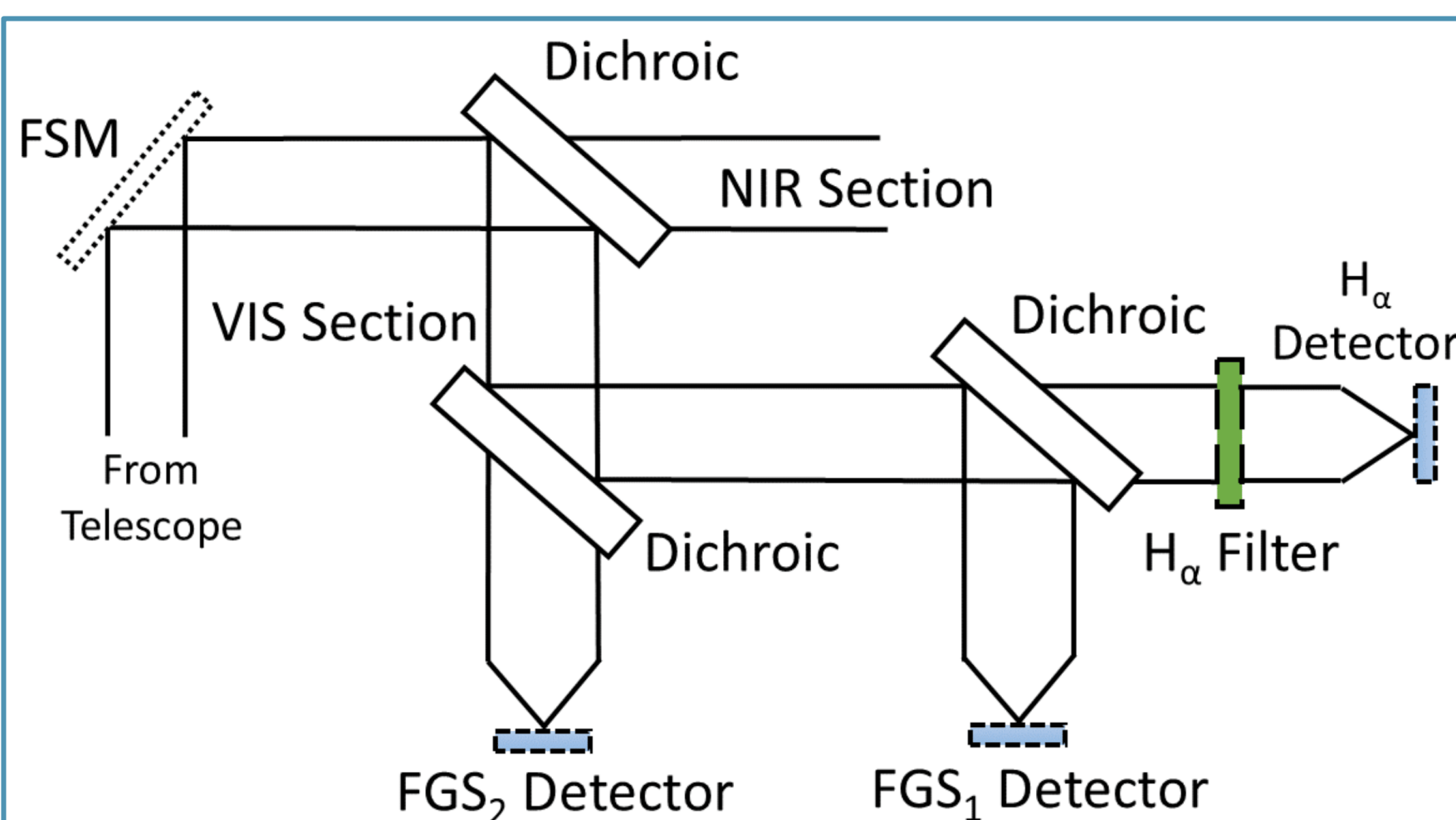
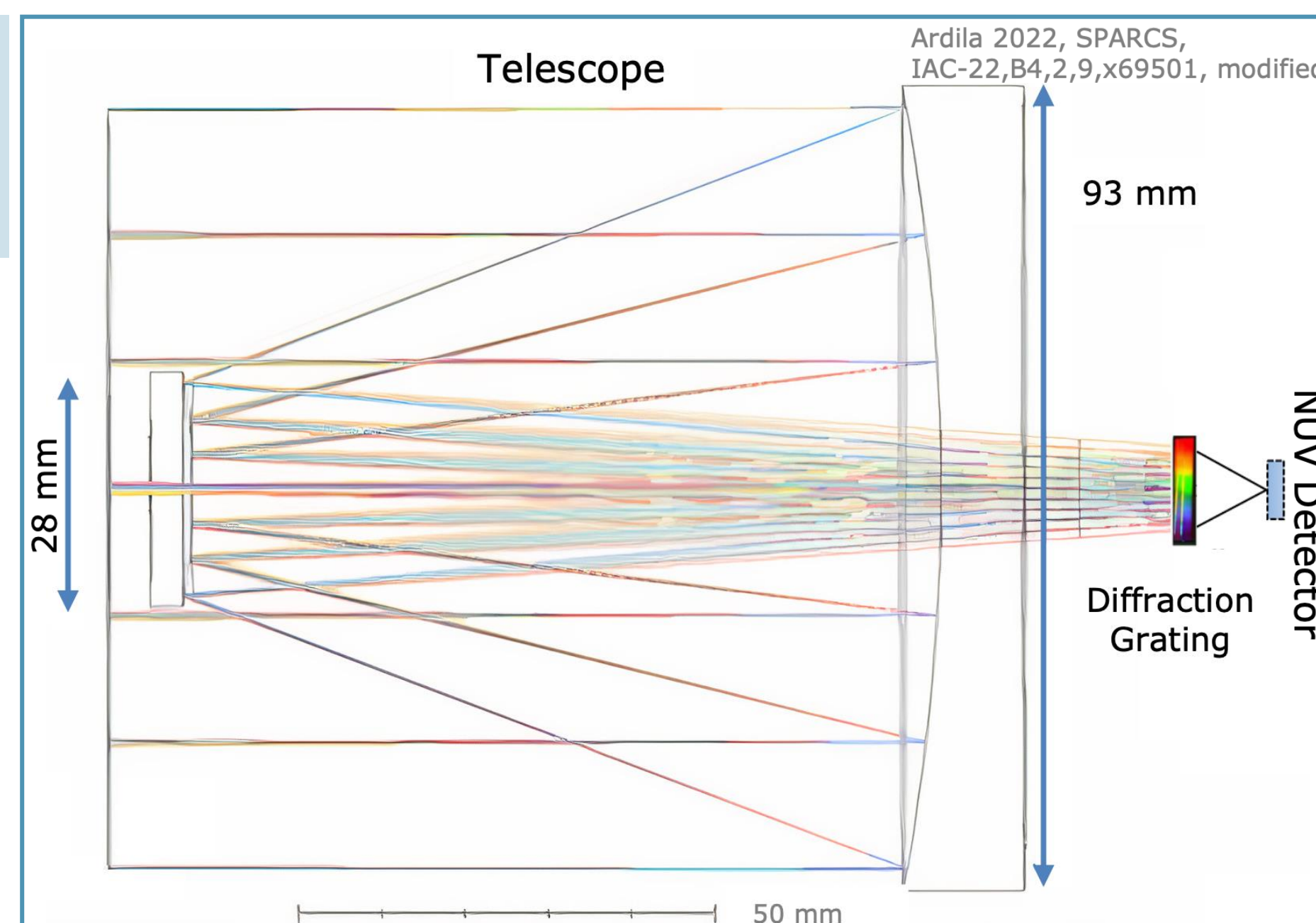
Payload

UV monitoring of stellar activity

- Ritchey-Chrétien configuration
- Primary mirror diameter: 9.3 cm
- FOV: 40 arcmin

Figure 1: Optical path and configuration of the UV telescope based on the near-ultraviolet part of the SPARCS CubeSat mission payload [8].

UV target: Mg II H&K, Ca II H&K lines (stellar activity tracers).



NIR observations of the planet

- Cassegrain telescope
- Main mirror diameter: 2.4 m
- FOV: 10 arcsec
- Focal length: 12.6 m

Figure 2: VIS lightpath of the VIS-NIR telescope. The fast-steering mirror increases telescope pointing precision through fast tilt-shift corrections on the collimated beam.

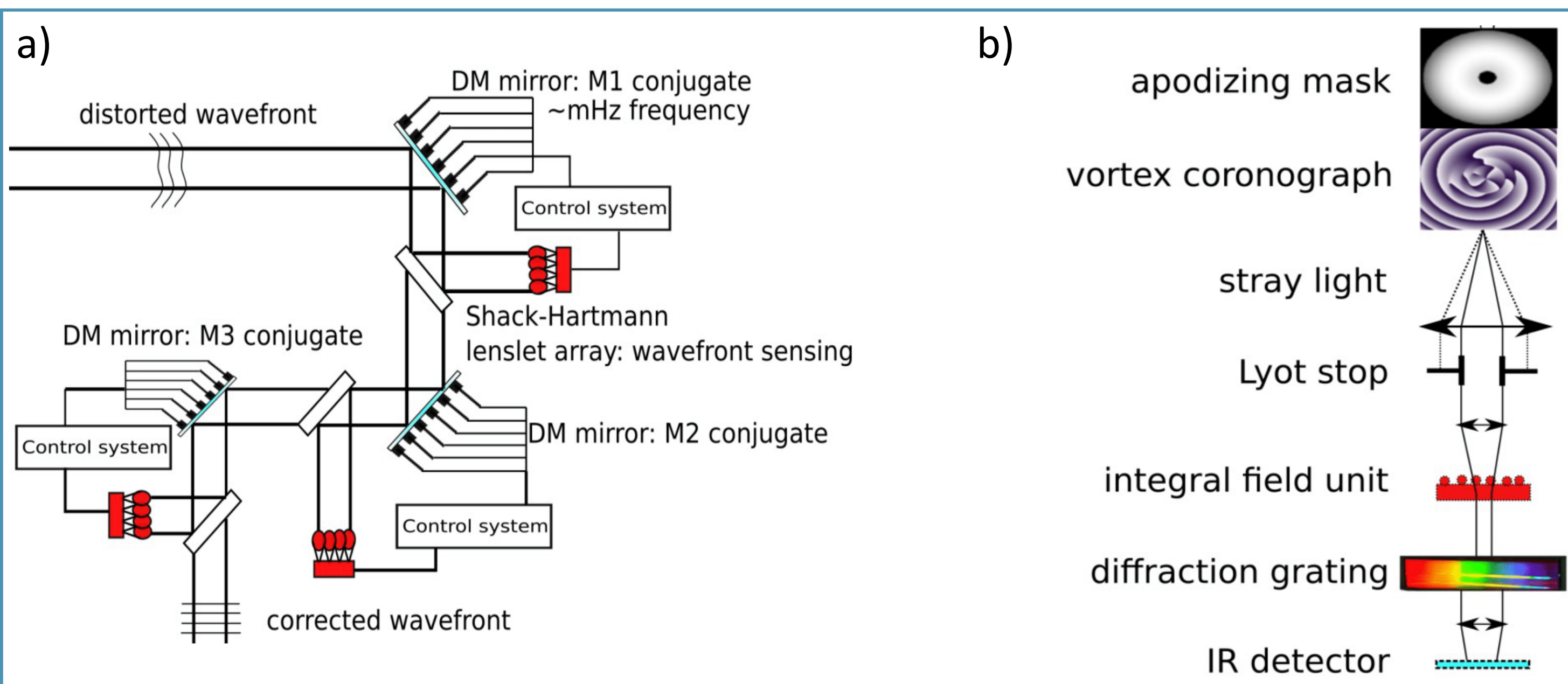


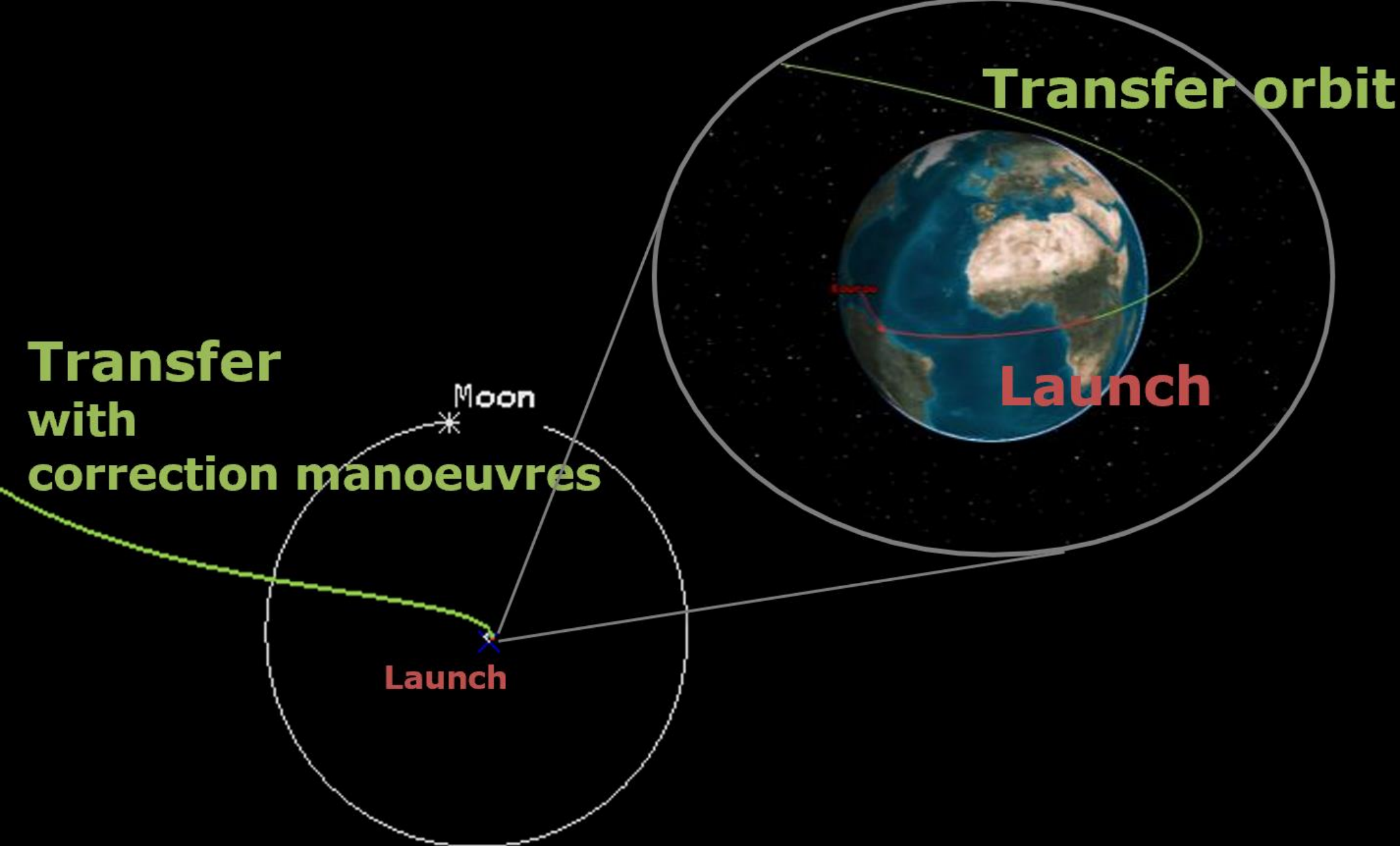
Figure 3: NIR lightpath of the VIS-NIR telescope. a) Adaptive optics system. b) MARY instrument.

NIR target: He 1083 nm line of planet's atmosphere (coronagraphy to block out star light)

VIS target: H α 656 nm line (trace stellar activity) + pointing corrections

Mission profile

Disposal



L2 Halo with orbit maintenance

Manoeuvre (with margins)	ΔV [m/s]
Launch + transfer	55.00
Orbit maintenance (5 years)	44.63
Disposal	15.75
Total	115.38

Spacecraft bus design

Overall dimensions:

- Length: 8.79 m
- Maximum width: 4.04 m
- Dry mass: 1599 kg
- Wet mass: 1783 kg

Main spacecraft bus design drivers:

- Instrument constraint: < 73 K
- Observation pointing precision: 50 mas performance drift error across 72 hours

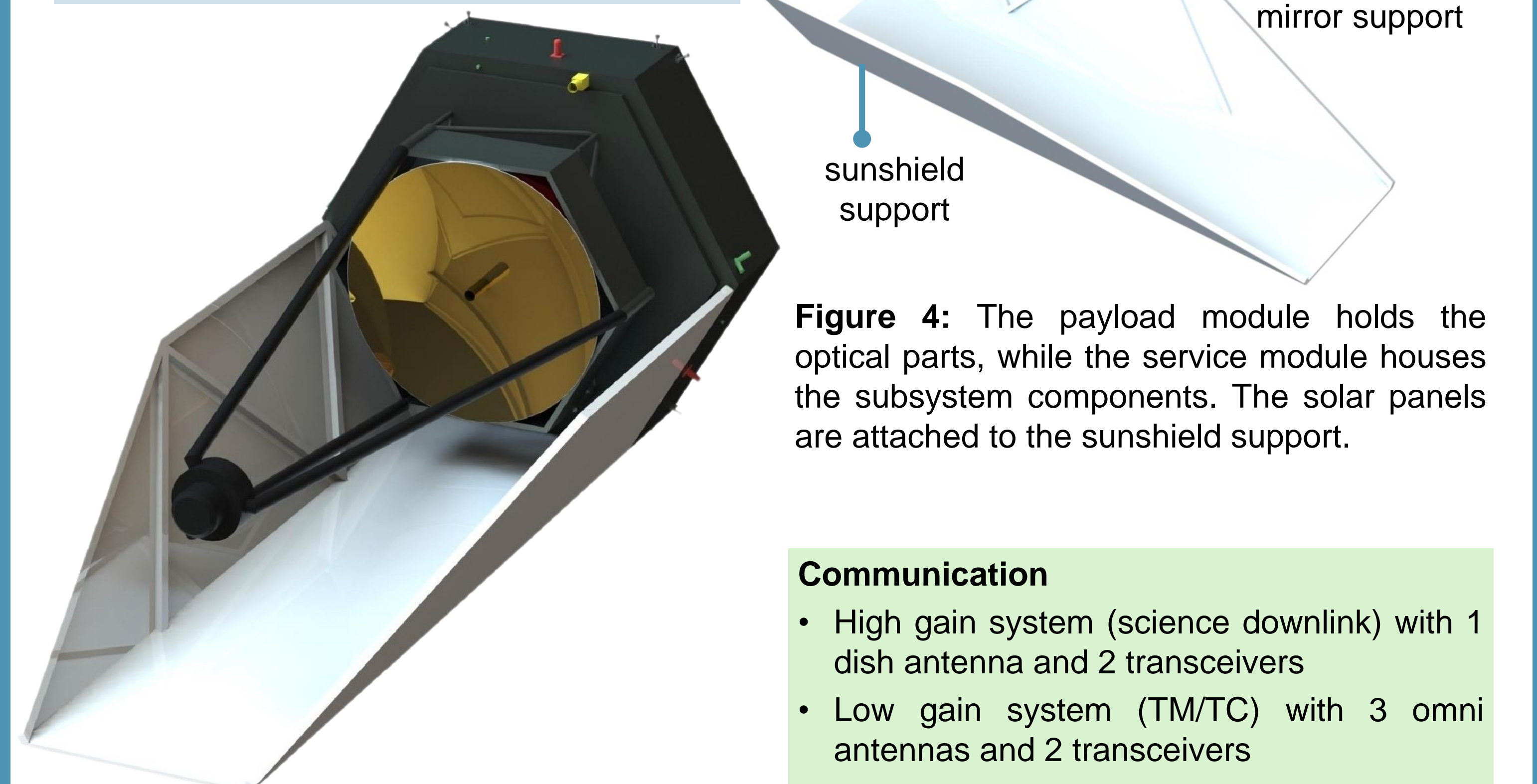


Figure 5: Overview of the Exodus spacecraft.

Propulsion

- 51 N thruster dimensioned for transfer burn into L2 orbit. (3 total for redundancy)
- Propellant: Monopropellant + hydrazine

ADCS

- Rough pointing: 6 sun sensors, gyroscope, 2 star trackers, 8 thrusters, 4 reaction wheels
- Fine pointing: Fine guidance sensor, fast steering mirror

Communication

- High gain system (science downlink) with 1 dish antenna and 2 transceivers
- Low gain system (TM/TC) with 3 omni antennas and 2 transceivers
- Data volume: up to 250 Gbits/week

Thermal

- 7-layer sunshield (213 K to 77.5 K)
- Thermal decoupling with heaters and radiators
- Surface coatings and insulators

Power

- Solar panel area: 5.72 m² (no eclipses during lifetime, driven by manoeuvre mode)
- Battery capacity: 2300 Wh (driven by launch mode)

Discussion and Conclusions

Iterations with the COMET tool at ESA Education Training Centre Concurrent Design Facility led to refinements of Exodus' subsystems.

- **Payload:** from a complex, low TRL single telescope design to two specialized telescopes — one for VIS-NIR observations and one for the UV spectral range.
- **FOV of MARY instrument:** changed from 3 to 10 arcsec, increasing the diversity and number of exoplanets observed.
- Simultaneous **observations of stellar activity and atmospheric loss** will reveal the impact of host stars on planetary systems.
- **High contrast** observations will **reduce technique biases** in new detections and enable comparisons with Solar System architecture.

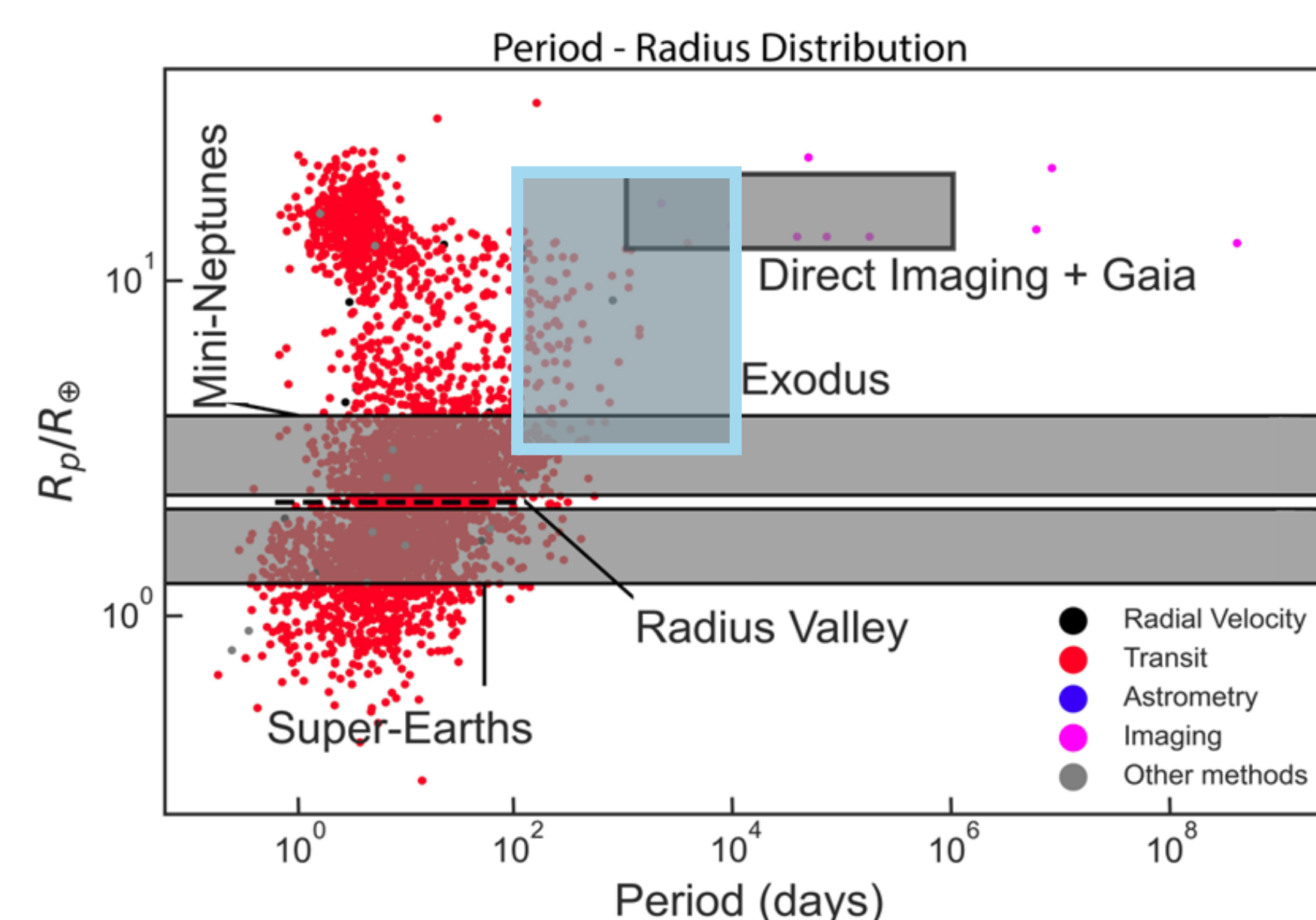


Figure 6: Exodus aims to monitor atmospheric escape in relation to stellar activity for over 5000 exoplanets, refining our understanding of exoplanet diversity and the conditions required for habitability.



Figure 7: Exodux spacecraft placed in Ariane 62 fairing.

References

- [1] B. J. Fulton et al., The California-Kepler Survey. III. A Gap in the Radius Distribution of Small Planets, 2017.
- [2] J. A. Johnson et al., Giant Planet Occurrence in the Stellar Mass-Metallicity Plane, 2010.
- [3] J. E. Owen, H. E. Schlichting, Mapping out the parameter space for photo evaporation and core-powered mass-loss, 2024.
- [4] A. Oklopčić, C. M. Hirata, A New Window into Escaping Exoplanet Atmospheres: 10830 Å Line of Helium, 2018.
- [5] J. J. Spake et al., Helium in the eroding atmosphere of an exoplanet, 2018.
- [6] A. M. Lagrange et al., A probable giant planet imaged in the β Pictoris disk. VLT/NaCo deep L'-band imaging, 2009.
- [7] G. Chauvin, Direct imaging of exoplanets: Legacy and prospects, 2024.
- [8] D. R. Ardila et al., The Star-Planet Activity Research Cube-Sat (SPARCS): Determining Inputs to Planetary Habitability, 2022.

Acknowledgements

The authors acknowledge funding from ESA and FFG, which supported the Alpbach Summer School 2023 and the Post-Alpbach Summer School Event 2023. The authors thank the tutors of the summer school, Günter Kargl and Leonard Schulz, for the valuable discussions. This study benefited from the work of 50 other students during the above-mentioned summer school and post-summer school event. The authors also thank all the tutors and organisers of the summer school for their support.

